

(2)

1964, Kanyagina, Z.V. & Zharitonov, A.V.
 "Soviet Astronomy AJ", Vol. 7, No. 6,
 P. 857. "Determination of the
 Stellar Magnitude of the Sun in
 Tricolor Systems based on Absolute
 Spectrophotometric measurements."

$$m_{IPV} = -26^m.79 \pm 0.02$$

$$m_V = -26.91 \pm 0.01$$

$$m_B = -26.27 \pm 0.01$$

16 stars

$$m_B - m_V = 0^m.64$$

Most. bright

Most. bright			V		B	
IPV						
m_*	star	$m_0 - m_*$	m_*	$m_0 - m_*$	m_*	$m_0 - m_*$
		$m_0 = -26.73$		$m_0 = -26.94$		$m_0 = -26.28$
0.34	β Ori	-27.07	0.15	-27.09	0.11	-26.39
		-26.75		-26.88		-26.29
0.14	α Lyr	-26.89	0.04	-26.92	0.04	-26.25
		-25.86		-26.86	1.02	-27.24
0.89	α Aql	-27.65	0.80	-27.66	1.25	-27.54
		-26.95		-28.31		-26.29
1.34	α Leo	-28.29	1.36	-26.95		
Promedio XXXXXX			-26.91		-26.26	
$m_0 = -26.57$						

IPV $\equiv$ international Photo visual